# DARWIN: a next-generation liquid xenon observatory for dark matter and DARWIN neutrino physics

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### WIMPs Dark matter

- Exposure:  $200 (t \times y)$
- 99.98% ER rejection (30% NR acceptance)
- Combined (S1+S2) energy scale

### DARWIN

DARWIN is a next-generation astroparticle observatory, based on a 2.6 m height by 2.6 m diameter dual-phase time projection chamber. With 50 ton of liquid (e) xenon (40 ton active), it will probe the WIMP phase space, down to the irreducible

### Neutrinoless double beta decay

- <sup>136</sup>Xe abundance of 8.9% in natural Xe.
- $\rightarrow$  3.5 t of active <sup>136</sup>Xe
- **Q-value = 2.458 MeV**

- Energy window 5-35  $keV_{NR}$
- Light yield: 8 PE/keV
- Also search for: ALPs, solar axions, dark photons



Spin Independent



Spin Dependent

### Solar neutrinos



neutrino background. But DARWIN is much more than just

a dark matter detector.



Dominating intrinsic backgrounds: <sup>222</sup>Rn,
 <sup>137</sup>Xe, 2νββ decays of <sup>136</sup>Xe and solar <sup>8</sup>B
 neutrinos.





Preliminary sensitivity at 90% C.L.  $\begin{array}{rcl}
12 \text{ txy} & \rightarrow & T_{1/2} > 1.3 \times 10^{27} \text{ yr} \\
60 \text{ txy} & \rightarrow & T_{1/2} > 3.0 \times 10^{27} \text{ yr}
\end{array}$ 



<sup>8</sup>B solar neutrinos









- Measure pp-neutrinos to test main energy production models of the Sun
- $v_X + e \rightarrow v_X + e$
- 361 events/ $(t \times y)$
- Also measurements of:
- $P(\nu_e \rightarrow \nu_e)$
- $\boldsymbol{\theta}_W$



## Current R&D

DARWIN full-(x,y) scale demonstrator Test components at real

diameter under real conditions flatness of electrodes strength of the extraction field x-y homogeneity of the drift field

### The Collaboration





Supernova neutrinos

Irreducible background  $\rightarrow$  Physics Channel •  $\nu + A \rightarrow \nu + A$ , all flavors

- Solar  $v_{^8B}$ :  $E_{th} > 1 \ keV_{nr}$ 
  - $\rightarrow$  90 events/( $t \times y$ )
  - SN burst detection up to
     65 kpc from Earth (5σ)

700 events from a 27 M<sub>☉</sub> SN at 10 kpc

- 29 research
  - groups 12 countries





### Photosensors R&D

Several groups are testing novel photosensors as replacements for PMTs, such as SiPMs, LHM and the ABALONE technology.



The DARWIN Collaboration, Zürich, 12.2018

Working towards a CDR and a TDR
DARWIN is part of the APPEC roadmap

